Problem 6.32. Consider a classical particle moving in a one-dimensional potential well u(x), as shown in Figure 6.10. The particle is in thermal equilibrium with a reservoir at temperature T, so the probabilities of its various states are determined by Boltzmann statistics.

(a) Show that the average position of the particle is given by

$$\overline{x} = \frac{\int x e^{-\beta u(x)} \, dx}{\int e^{-\beta u(x)} \, dx},$$

where each integral is over the entire x axis.

Figure 6.10. A one-dimensional potential well. The higher the temperature, the farther the particle will stray from the equilibrium point.



(b) If the temperature is reasonably low (but still high enough for classical mechanics to apply), the particle will spend most of its time near the bottom of the potential well. In that case we can expand u(x) in a Taylor series about the equilibrium point x_0 :

$$u(x) = u(x_0) + (x - x_0) \frac{du}{dx}\Big|_{x_0} + \frac{1}{2}(x - x_0)^2 \frac{d^2u}{dx^2}\Big|_{x_0} + \frac{1}{3!}(x - x_0)^3 \frac{d^3u}{dx^3}\Big|_{x_0} + \cdots$$

Show that the linear term must be zero, and that truncating the series after the quadratic term results in the trivial prediction $\overline{x} = x_0$.

Phys 335: Problem 6.32: Thermal Expansion How can we relate the partition function to measurable quantities, such as there expansion ? (a) Average position: Motion in a potential well ubx) $\overline{\chi} = \int_{-\infty}^{\infty} \chi e^{-\beta u(x)} d\chi$ $\int_{-\beta u(x)}^{\infty} dx$ (5) Sample U(x) U(x) No Why does in creasing Timeseure x? Why doem't TX Stay the same, even if oscillations get bigger? The will is a-symmetric. Taylo Series $\begin{array}{rcl}
\mu(\chi) = & \mu(\chi_{0}) + \begin{pmatrix} d \upsilon \\ d \chi \end{pmatrix}_{\chi_{0}} & \chi_{0} \end{pmatrix} + \frac{1}{2!} \frac{d^{2} \upsilon}{d \chi^{2}} & (\chi - \chi_{0})^{2} \\
+ & \frac{1}{3!} \frac{d^{3} \mu}{d \chi^{3}} \begin{pmatrix} \chi - \chi_{0} \end{pmatrix}^{3} + \cdots & 3pring \ constant \end{array}$ At the minimum, du/=0 $u(x) = u(y_{\delta}) + a(x-y_{\delta})^{2} + b(x-y_{\delta})^{3}$ what $\overline{u}(\overline{x})^{2}$

- 2 -First, suppre we only take the first 2 terms $\overline{\chi} = \int \chi e^{-\beta u(\chi)} d\chi$ $\int e^{-\beta u(n)} dy$ Numeration $\int x e^{-\beta u_0} -\beta a(x-x_0)^2 dx$ Let Z= x- x = > x= Z+X, dx=dZ $N = \int (Z + \eta_0) e^{-\beta u_0} e^{-\beta a z^2} dz$ $= \int Ze^{-\beta u_{o}} -\beta az^{2} + \chi_{o}e^{-\beta u_{o}} \int e^{-\beta az^{2}} dz$ Denominator: $D = \int_{-\infty}^{\infty} e^{-\beta u l_x} dx = \int_{-\infty}^{\infty} e^{-\beta u l_x - x_0^2} dx$ $= \int e^{-\beta \mu_0} - \beta a z^2 dz$ $\overline{\mathcal{R}} = \mathcal{N}$ \mathcal{N} \mathcal{N} \mathcal{N} \mathcal{N}

- (c) If we keep the cubic term in the Taylor series as well, the integrals in the formula for \overline{x} become difficult. To simplify them, assume that the cubic term is small, so its exponential can be expanded in a Taylor series (leaving the quadratic term in the exponent). Keeping only the largest temperature-dependent term, show that in this limit \overline{x} differs from x_0 by a term proportional to kT. Express the coefficient of this term in terms of the coefficients of the Taylor series for u(x).
- (d) The interaction of noble gas atoms can be modeled using the Lennard-Jones potential,

$$u(x) = u_0 \left[\left(\frac{x_0}{x}\right)^{12} - 2\left(\frac{x_0}{x}\right)^6 \right].$$

Sketch this function, and show that the minimum of the potential well is at $x = x_0$, with depth u_0 . For argon, $x_0 = 3.9$ Å and $u_0 = 0.010$ eV. Expand the Lennard-Jones potential in a Taylor series about the equilibrium point, and use the result of part (c) to predict the linear thermal expansion coefficient (see Problem 1.8) of a noble gas crystal in terms of u_0 . Evaluate the result numerically for argon, and compare to the measured value $\alpha = 0.0007 \text{ K}^{-1}$ (at 80 K).

D What if you include the cubic? $N = \int (2+\gamma_0) e^{-\beta u_0} e^{-\beta az^2} e^{-\beta bz^3}$ what to do here? No elsed solution Approximate! Assure $\beta bz^3 << 1$ (i.e. extra energy $bz^3 << kT$) (c)d Z

Then $e^{-\beta bz^3} \approx 1 - \beta bz^3$ $N = \int (z + \gamma_s) e^{-\beta u_s} e^{-\beta az^2} (1 - \beta bz^3) dz$ 4 tame. Loch at symmetries to see what is O $N = \int Z \cdot 1 e^{-\beta u_0} e^{-\beta a z^2} d z +$ $\int \gamma_0 \cdot 1 e^{-\beta u_0 - \beta a z^2} dz$ $\int -\beta b z^4 e^{-\beta u_0} -\beta a z^2 dz$ $\int_{-\infty}^{\infty} -\mathcal{N}_{o}\left(\beta b z^{3}\right) e^{-\beta u \cdot o} e^{-\beta a z^{2}} z$ N= No Soo e-Bus -Baz dz - Bb Szy e-Bus -Baz dz Devonin ator: Mole the same appriximation $D^2 \int_{a}^{a} e^{-\beta u_0} e^{-\beta a z^2} (1 - \beta b z^3) dz$ integrates to 0 R= No - Bbe-Buo 24e-Baz dz $e^{-\beta u_0} \int e^{-\beta a z^2} dz$ How to evaluate such integrals ?

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Step 1 = Make deview storless l.g M = JBa Z dy = JBa dz = dz = JBa dy $\int_{-\infty}^{\infty} e^{-\beta a z^2} dz = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} e^{-y^2} dy$ JTT See Appendix B.1. Numerator: See Appendix B.1, esp. publer B.2 $\int_{-\infty}^{\infty} z^{\gamma} e^{-\beta \alpha z^{2}} dz = \frac{1}{(\beta \alpha)^{2}} \frac{1}{\sqrt{\beta \alpha}} \int_{-\infty}^{\infty} y^{\gamma} e^{-\gamma} dy$ This is 3/TT $\overline{X} = X_0 - \beta \overline{b} \cdot \frac{1}{(\beta_a)^2} \cdot \frac{1}{\sqrt{\beta_a}} \cdot \frac{3}{4} \sqrt{\pi}$ JBa NTT $\overline{\chi} = \chi_0 - \frac{3}{4} \cdot \frac{b}{Ba^2} = \chi_0 - \frac{3}{4} \cdot \frac{b}{Ba^2} kT$ Units check: Recall $u(r) = u_0 + a(r-r_0)^2 + b(r-r_0)^3$ $\begin{bmatrix} \frac{b}{\beta a^2} \end{bmatrix} = \frac{J/m^3}{-1} \cdot (J/m^2)^2 = m V$

what is the sign for b? Depends on shape of pitatiol. $b = \frac{1}{31} \frac{d^3 u}{d x^3}$ The the megative if it gets less steep X > x o due to themal expansion. Example: Lennard-Jones Potential $u(x) = u_{o}\left[\frac{x_{o}}{x}\right]^{R} = 2\left(\frac{x_{o}}{x}\right)^{6}$ Is this familian? where does it come from ? Argon: $\chi_0 = 0.39 \text{ mm}$ $U_0 = 0.010 \text{ eV}$ $\sum \chi_1 = 0.0807 \text{ f}$ at 80 K. L tem: Dipole Dipole inderaction An Fluctuation field depole x3 And induced dipole interaction energy $-\frac{p^2}{\chi^3}$ Short range repulsion complex / exclusion principle as altitals overlaps, Had to calculate. Easy numerous approximation: You already have 1/26 Square it to get ______

More on the dipole interaction \vec{P} , \vec{b} (instantaneous) depole moment \vec{b} \vec{E} , due to \vec{P} , $\propto \frac{1}{N^3}$ E, due to P, Induces dipolo P2 x E,. Interaction energy is (Eq. 4-26 in Jackson) (for 2 parallel dipole moments) $U_{12} = -\frac{2p_1p_2}{4\pi\epsilon_0 R^3}$ and since $p_2 \propto \frac{1}{\gamma^3}$ U,2 x -1 x6

$$-\frac{1}{16} = \frac{1}{16} \left[\frac{1}{16} \frac{1}$$

Then $\overline{\chi} = \chi_0 - \frac{3}{4} \frac{b}{RZ} kT$ $\overline{\mathcal{X}} = \mathcal{X}_{6} - \frac{3}{4} \left(\frac{-252 \, u_{0} / \chi_{0}^{3}}{(36 \, u_{0} / \chi_{0}^{2})^{2}} \right) k T$ $= \mathcal{Y}_{6} \left[1 + \frac{3 \cdot 252 \, k T}{4 \, (36)^{2} \, u_{0}} \right]$ = N, II + XT $\begin{aligned} \chi &= \frac{3}{4} \frac{252}{(36)^2} \frac{8.617 \times 10^{-5} \text{eV}/\text{k}}{0.00126/\text{k}} = \frac{0.00126}{/\text{k}} \\ \chi_{exp} &= 0, \ 0.007 \ /\text{k} \end{aligned}$ So we get to within a power of 2. Take - aways 1. Boltzmann factor can lead to real physical measure able quantities 2. Handling integrals + calculations - Scaling - approximations