Problem 6.32. Consider a classical particle moving in a one-dimensional potential well $u(x)$, as shown in Figure 6.10. The particle is in thermal equilibrium with a reservoir at temperature T , so the probabilities of its various states are determined by Boltzmann statistics.

(a) Show that the average position of the particle is given by

$$
\overline{x} = \frac{\int xe^{-\beta u(x)} dx}{\int e^{-\beta u(x)} dx},
$$

where each integral is over the entire x axis.

Figure 6.10. A one-dimensional potential well. The higher the temperature, the farther the particle will stray from the equilibrium point.

(b) If the temperature is reasonably low (but still high enough for classical mechanics to apply), the particle will spend most of its time near the bottom of the potential well. In that case we can expand $u(x)$ in a Taylor series about the equilibrium point x_0 :

$$
u(x) = u(x_0) + (x - x_0) \frac{du}{dx}\bigg|_{x_0} + \frac{1}{2}(x - x_0)^2 \frac{d^2u}{dx^2}\bigg|_{x_0} + \frac{1}{3!}(x - x_0)^3 \frac{d^3u}{dx^3}\bigg|_{x_0} + \cdots
$$

Show that the linear term must be zero, and that truncating the series after the quadratic term results in the trivial prediction $\bar{x} = x_0$.

Phys 335: Problem 6.32: Thermal ExpansionHow can we relate the partition function to measurable quantities, such as their expansion (a) Average position: Motion in a prtential well $u|\chi$) \overline{X} = $\int_{-\infty}^{\infty}$ χ $e^{-\beta u l \phi}$ dy $\int_{0}^{\infty} e^{-\beta u(\tau)} d\tau$ b) Sauple Ulx, $\mathcal{U}(\mathcal{X})$ π L Why does in creasing Timerease X . Why doem $\frac{1}{X}$ Stay the Sane, even if oscillations get bigger! The <u>Taylo Series</u> $u(\gamma) = u(\gamma_0) + \left(\frac{dU}{d\gamma}\right)_{\gamma_0} \frac{(\gamma_-\gamma_0) + \frac{1}{2!} \frac{dU}{d\gamma^2}}{u^2}$ (x-1) $\frac{1}{1} \frac{d^3u}{d\gamma^3}\bigg|_{\gamma_5} (\chi \gamma_5)^3 + \cdots$ Spring constant At the minimum, du $\begin{array}{c} d \circ \\ d \times \alpha_6 \end{array}$ $u \left(\gamma \right) = u \left(\gamma_{0} \right) + a \left(\gamma - \gamma_{0} \right)^{2} + b \left(\gamma - \gamma_{0} \right)^{3}$
what $\overline{\nu}$ $\overline{\gamma}$?

Kirst, suppue we only take the first 2 terms $\pi = \frac{\int \chi e^{-\beta u(x)} dx}{\int e^{-\beta u(x)} dx}$ Nuveration $\int \pi e^{-\beta u_{0} - \beta a (x-x)} dx$ Let $Z = x - \gamma_0 \implies x = Z + \gamma_0$, dx = dz $N = \int (Z + \alpha) e^{-\beta u} e^{-\beta a z^2} dz$ $=$ $\int Ze^{-\beta u_{0}-\beta az^{2}} dx + \lambda_{0}e^{-\beta u_{0}} \int e^{-\beta az^{2}} dz$ Denominator. On Cald Func. avec $D = \int_{0}^{\infty} e^{-\beta u/x} dx = \int_{0}^{\infty} e^{-\beta u} e^{-\beta (x-x)^{2}} dx$ $=\int_{0}^{\infty}e^{-\beta\mu_{0}}e^{-\beta_{a}z^{2}}dz$ 2 = N = N = Mm unsurprising.

- (c) If we keep the cubic term in the Taylor series as well, the integrals in the formula for \bar{x} become difficult. To simplify them, assume that the cubic term is small, so its exponential can be expanded in a Taylor series (leaving the quadratic term in the exponent). Keeping only the largest temperature-dependent term, show that in this limit \bar{x} differs from x_0 by a term proportional to kT . Express the coefficient of this term in terms of the coefficients of the Taylor series for $u(x)$.
- (d) The interaction of noble gas atoms can be modeled using the Lennard-Jones potential,

$$
u(x) = u_0 \left[\left(\frac{x_0}{x} \right)^{12} - 2 \left(\frac{x_0}{x} \right)^6 \right].
$$

Sketch this function, and show that the minimum of the potential well is at $x = x_0$, with depth u_0 . For argon, $x_0 = 3.9$ Å and $u_0 = 0.010$ eV. Expand the Lennard-Jones potential in a Taylor series about the equilibrium point, and use the result of part (c) to predict the linear thermal expansion coefficient (see Problem 1.8) of a noble gas crystal in terms of u_0 . Evaluate the result numerically for argon, and compare to the measured value $\alpha = 0.0007 \text{ K}^{-1}$ (at 80 K).

 $\begin{array}{rcl}\nD & \text{other than} & \text{other than}$ (c) d z What to do here? No closed solutive
Approximate! Assure Bbz³ << 1
(i.e. extra energy bz³ << RT)

Then $e^{-\beta b z^{3}} \approx 1 - \beta b z^{3}$ $N = \int_{0}^{\infty} (z+y) e^{-\beta u} e^{-\beta a z} (1-\beta b z^{3}) dz$ 4 terme. Look at symmetries to see what is 0 $N = \int Z \cdot 1 \ e^{-\beta u_0} e^{-\beta a z^2} \ Q \ z^2$ \int_{0}^{∞} χ_{0} 1 e βu_{0} $\beta a t^{2}$ d t \int_{0}^{∞} $- \beta b z^4 e^{-\beta u_v} e^{-\beta a z^2} dz$ $\int_{-\infty}^{\infty} -\gamma_{0}(\beta b z^{3}) e^{-\beta u_{0}}e^{-\beta a z^{2}}dz$ $N = \gamma_0 \int_{\infty}^{\infty} e^{-\beta u_s} e^{-\beta a z^2} dz - \beta b \int z^y e^{-\beta u_s} e^{\frac{z^2}{2}} dz$ Denominator Make the Sane approximation $D = \int_{-\infty}^{\infty} e^{-\beta^{u_{b}}} e^{-\beta a \tau} \left(1 - \beta b \tau^{5}\right) d\tau$ Integrates to ⁰ $\widehat{\pi}$ = μ_{0} - β b $e^{-\beta u_{0}}$ $\int_{0}^{\infty} z^{4} e^{-\beta a z^{2}} dz$ $e^{-\beta u_{o}}\int_{0}^{\infty}e^{-\beta a\tau^{2}}d\tau$ How to evaluate such mtegrals?

Step 1: Make deviensronless lg $y = \sqrt{\beta \alpha^2}$ 2 $dy = \sqrt{\beta a} dz \implies dz = \frac{1}{\sqrt{\beta a}} dy$ $\int_{-\infty}^{\infty} e^{-\beta a z^{2}} dz = \frac{1}{\sqrt{\beta a}} \int_{-\infty}^{\infty} e^{-\frac{y^{2}}{a}} dy$ \int $\frac{1}{\sqrt{\pi}}$ See Appendix B.1
Numerator: See Appendix B.1, esp problem B.2 $\int_{-\infty}^{\infty} z^{\frac{y}{2}} e^{-\beta a z^2} dz = \frac{1}{(\beta a)^2} \int_{\beta a}^{\infty} y^{\frac{y}{2}} e^{-y^2} dy$ This is $\frac{3}{4}\sqrt{\pi}$ $\frac{1}{\pi}$ = N_0 - βb . $\frac{1}{(\beta a)^2}$, $\frac{1}{\sqrt{3}a}$, $\frac{3}{4}$ $\frac{1}{\sqrt{\beta a}}$ $\sqrt{\pi}$ $\pi = x_{0} - 3$ $\frac{b}{4} - \frac{b}{8a^{2}} = x_{0} - \frac{3}{4} \frac{b}{a^{2}} kT$ Units check: Recall $u(r) = u_0 + a (r-r_0)^2 + b (r-r_0)^3$ $\left[\frac{b}{\beta a^{2}}\right] = \frac{J/m^{3}}{1-(J/m^{2})^{2}} = m$

 $-5-$

what is the sign for b ? Depends on shape of pritative. $b = \frac{1}{3!} \frac{d^3u}{d\chi^3}$ x_b is negative if it getslesssteep $x \rightarrow \infty$ due to themal expansion. Example Lennard Jones Potential $u(\gamma) = u_{0} \left(\frac{\gamma_{0}}{\gamma}\right)^{12} - 2\left(\frac{\gamma_{0}}{\gamma}\right)^{6}$ Is this familiar?
Where does it come from? $M_{0} = 0.37$ and $M_{0} = 0.010eV$ Expt $x = 0.0007/k$ at 80 k Le temme Dipole / Dipole interaction An)
Induced dipole, $\begin{array}{c|cc}\n\hline\n\text{(A)} & \text{(A)} \\
\hline\n\text{(A)} & \text{(B)} \\
\hline\n\text{(B)} & \text{(C)} \\
\text{(D)} & \text{(D)} \\
\hline\n\text{(E)} & \text{(E)} \\
\hline\n\text$ interaction energy Short range respulsion complex / exclusion principle as orbitals overlaps. Easy numerous approximation : You already have /x 6 Square it to get $\frac{1}{x^{\frac{1}{2}}}$

 $-7-$ More on the depole interaction Pi à (instantaneous) dépôle monent à E due to p, inducer dipole p & E,. Internetion energy is (Eg. 4-26 in Jackson) U_{12} = -2 p, p and since $P_{22} \propto \frac{1}{\gamma^{3}}$
4 T 6 g p 3 $U_{12} \propto -1$
 N^{6}

3
\n
$$
\frac{1}{16\pi} \int_{0}^{1} \frac{1}{\sqrt{12}} \
$$

Then $\overline{\gamma}$ = γ_0 - $\frac{3}{4}$ b k T $\frac{x_{6}-3}{4}\frac{-252u_{6}/x_{6}^{2}}{(36u_{6}/x_{6}^{2})^{2}}$
 \sim 5. 3. 252 kT γ $= \gamma_{0} \left[1 + \alpha T\right]$ $22 = 3$ ds2, 8.617 N/D eV/k 0.00126/k
4 (36)2 0.01 eV
21 20117/11 α_{exp} = 0, 0007 /K So we get to within a power of 2. Iahe-away)
1. Boltzmann factor can lead to real sphysical measure afte quantities 2 Handling integrals calculations Scaling approximations